

## PHYS 643 Computational Exercise 2: Steepening

You should email me a write-up of your results and the code that you used to do the calculation and make plots before the class on October 23.

*Overview.* The goal of this exercise is to write a 1D hydro code and use it to demonstrate the steepening of a sound wave.

*Algorithm.* The lecture notes on hydrodynamics from the University of Heidelberg give a simple algorithm that you can use to solve the 1D hydro equations (in Chapter 5). First, the fluid equations are written in flux-conservative form, with conserved quantities

$$f_1 = \rho$$

$$f_2 = \rho v$$

(the mass and momentum densities) and it is assumed that  $P = c_s^2 \rho$  with constant sound speed  $c_s$ . The equations to solve are then

$$\frac{\partial f_1}{\partial t} + \frac{\partial}{\partial x} (v f_1) = 0$$

$$\frac{\partial f_2}{\partial t} + \frac{\partial}{\partial x} (v f_2) = -\frac{\partial P}{\partial x}.$$

These are in flux-conservative form with the pressure gradient acting as a source term for the momentum density  $f_2$ . Note that given  $f_2$  and  $f_1$ , the velocity at the grid centre can be obtained from the ratio  $f_2/f_1$ .

The algorithm has two steps:

1. Use donor-cell advection to update  $f_1$  and  $f_2$ . To calculate the velocity at the cell boundaries, you can take an average of the velocity at the cell centres

$$v_{j+\frac{1}{2}} = \frac{1}{2} (v_j + v_{j+1}).$$

2. Add an additional update to the value of  $f_2$  from step 1 to take into account the source term. You can do this by writing

$$\frac{\partial P}{\partial x} = c_s^2 \frac{\partial \rho}{\partial x}$$

and using a first order difference to calculate the density gradient in terms of the new values of  $f_1$  you found in step 1.

### Questions

1. Choose an initial condition that has a sinusoidal variation in density and/or velocity. Check that for small amplitudes, the wave propagates as expected.

2. Do you see steepening at larger wave amplitudes?
3. How large a timestep can you take and still be numerically stable?
4. Do you form a shock in your simulation? What sets its thickness?

### Possible extensions

- An interesting extension is to add the energy equation to your code. We need a third quantity

$$f_3 = \rho e_{\text{tot}}$$

where  $e$  is the specific total energy (sum of kinetic and internal energy). The energy equation is

$$\frac{\partial f_3}{\partial t} + \frac{\partial}{\partial x} (v f_3) = -\frac{\partial}{\partial x} (vP).$$

As for momentum, we can solve this by advecting  $f_3$  in step 1 and then updating  $f_3$  using the source term on the right hand side in step 2. The pressure is

$$P = (\gamma - 1)\rho e$$

where

$$e = e_{\text{tot}} - \frac{v^2}{2}$$

is the thermal energy.

With the energy equation included, you will be able to validate the shock jump conditions, e.g. check that the compression factor is  $(\gamma + 1)/(\gamma - 1)$  for a strong shock.

- As a next step, you could modify your code to work in spherical symmetry (adding appropriate  $r^2$  factors), and then model the Sedov-Taylor blast wave.
- Another area to investigate is to use a higher order approximation for the advection step (section 4.3 in the Heidelberg notes) and see how it improves the results.