



EXOPLANET HUNTING 101

PHYS 182 - Guest Lecture

Sept 20, 2018 11:35-1:05 PM

BY LISA DANG



WHAT IS A PLANET?

International Astronomical Union Definition

- In orbit around the Sun
- Massive enough to be round
- Cleared its orbit of debris



WHAT IS A PLANET?

International Astronomical Union Definition

- In orbit around the Sun
- Massive enough to be round
- Cleared its orbit of debris



PLUTO



WHAT IS AN EXOPLANET?

Also referred as exosolar planets or extrasolar planet

According to the IAU

- In orbit around a star
- Not massive enough for fusion of deuterium



WHAT IS AN EXOPLANET?

Also referred as exosolar planets or extrasolar planet

According to Lisa Dang (et al.)

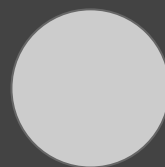
- In orbit around a star
- Not massive enough for fusion of deuterium
- + Rogue (free-floating) planets



IS OUR SUN SPECIAL?

Not really, it's a pretty average star.

- 10 % of all stars are Sun-like stars
- 20 billions star in our Galaxy are just like our Sun



BUT YET, IT'S DIFFERENT

SUN



- Was the only known planetary system for centuries
- Still the only system hosting 9 ± 1 significantly diverse planets
- We think we understand how they formed



Finding is HARD.

(but we've gotten better at it!)

1. **They are small and far.**
2. **They are dim compared to their host star.**

Sun



**1 IT ALL
STARTED
WITH
51 Peg b
(aka Dimidium)**



1 IT ALL
STARTED
WITH
51 Peg b
(aka Dimidium)





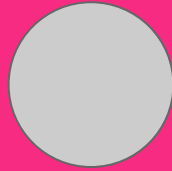
Hot Jupiter? What?

51 Pegasi b

- Mass of 0.5 Jupiter mass
- Takes 4.23 days to orbit around its star
- Temperature over 1300K (Venus is about 700K)




What's next?

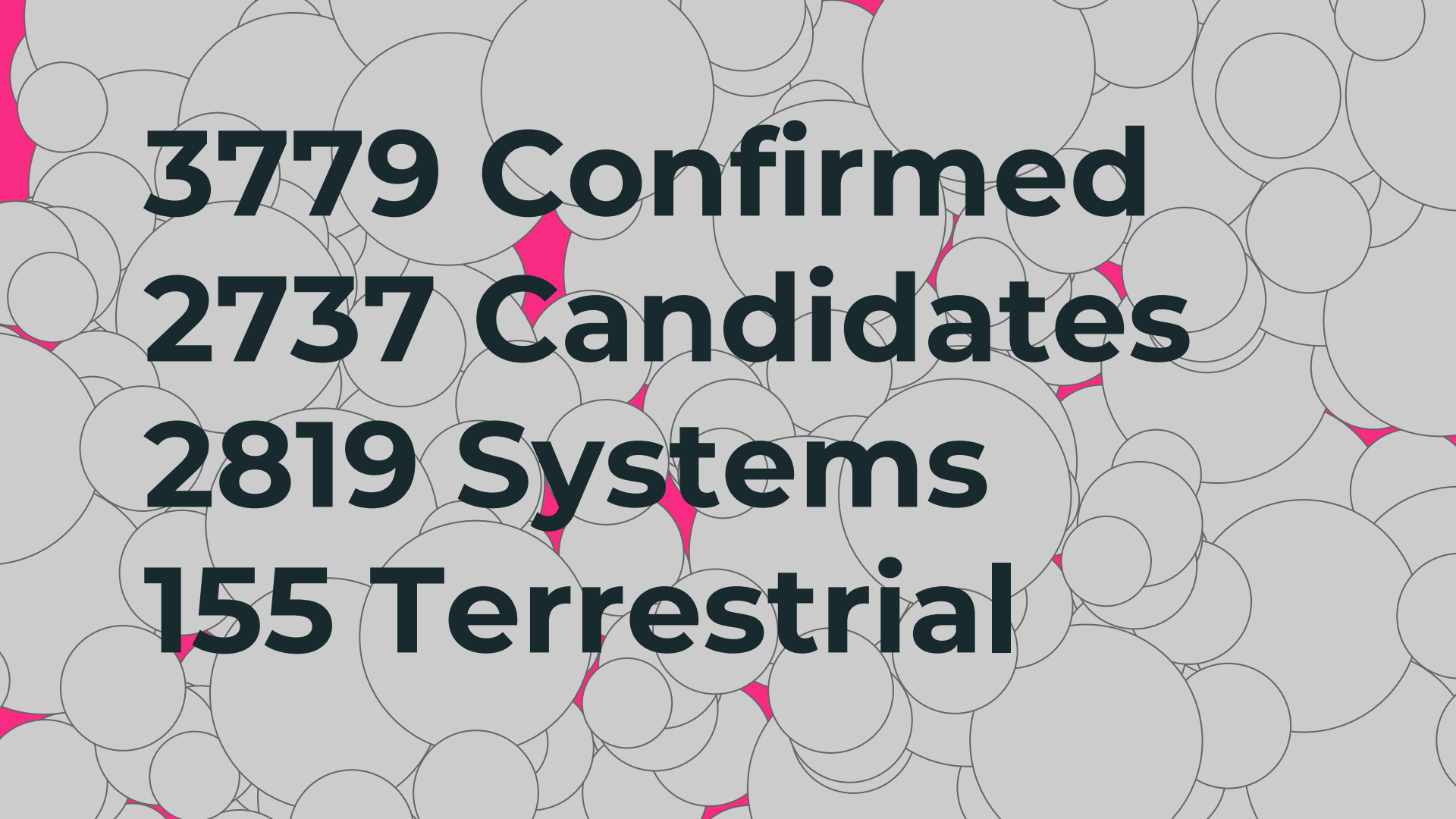


**FIND.
MORE.
EXOPLANETS.**



The background consists of a dense, overlapping pattern of circles. Most circles are light gray with thin black outlines. Interspersed among these are several bright pink circles, also with thin black outlines. The circles vary in size and are arranged in a way that creates a textured, bubbly effect.

**FIND.
MORE.
EXOPLANETS.**

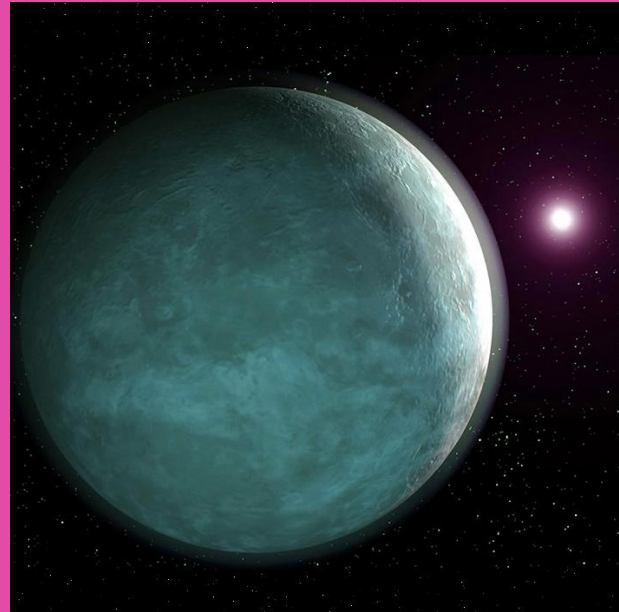


3779 Confirmed
2737 Candidates
2819 Systems
155 Terrestrial



Most recent planet discovery: Wolf 503 b

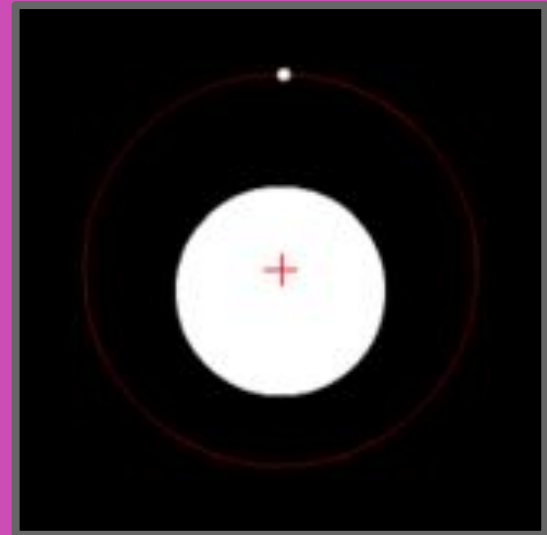
A Neptune-like
planet
**discovered by
student in
Montreal!**





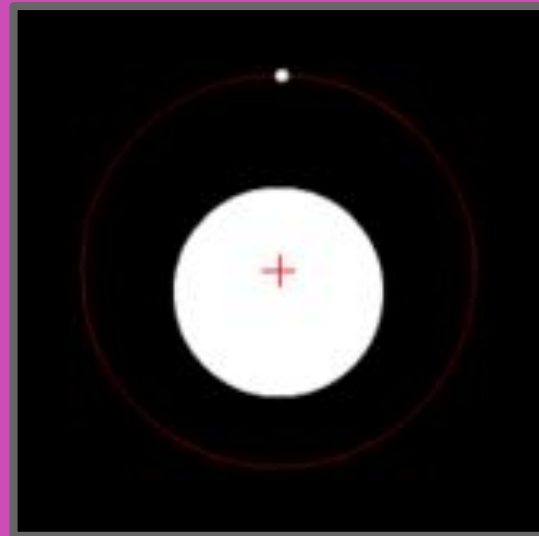
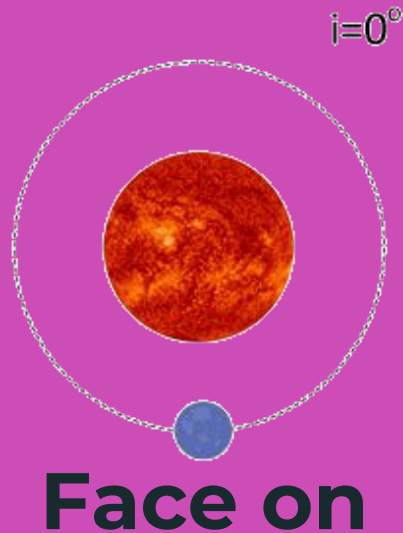
STARS' INVISIBLE DANCE PARTNER

In the presence of **another body mass** orbiting a star, the star will also orbit about their **common centre of mass**





STARS' INVISIBLE DANCE PARTNER



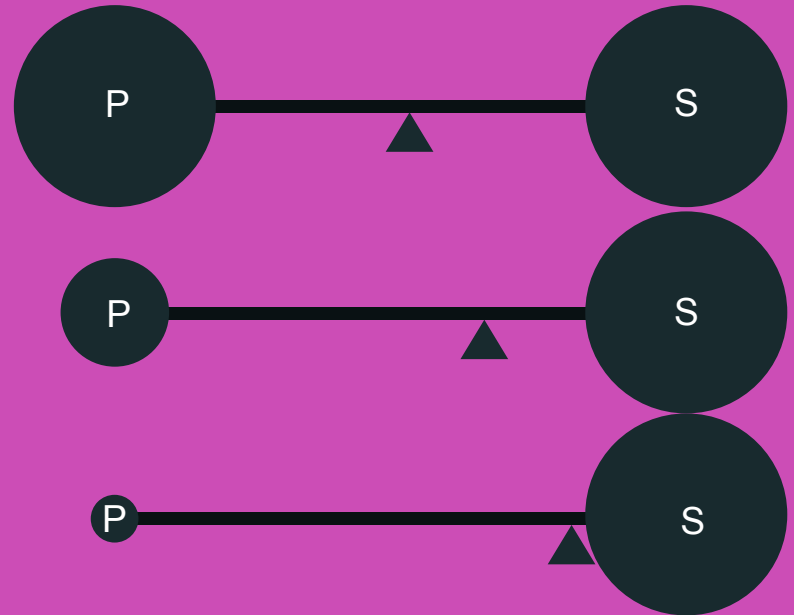


Centre of Mass

Definition:

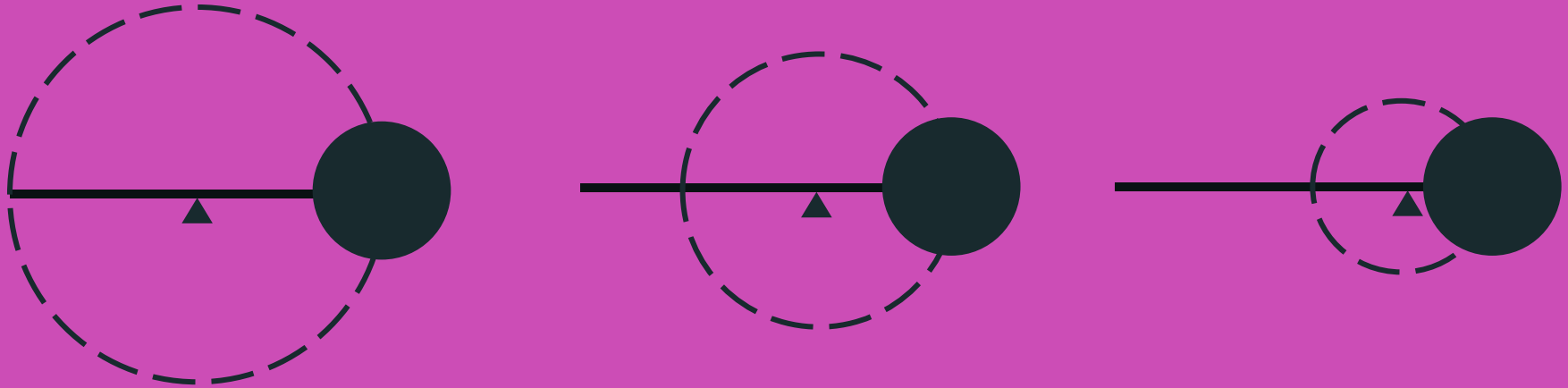
Mean location of distribution of mass in space.

*The **heavier** the star is compared to the planet, the **closer** the centre the center of mass is to the star.*





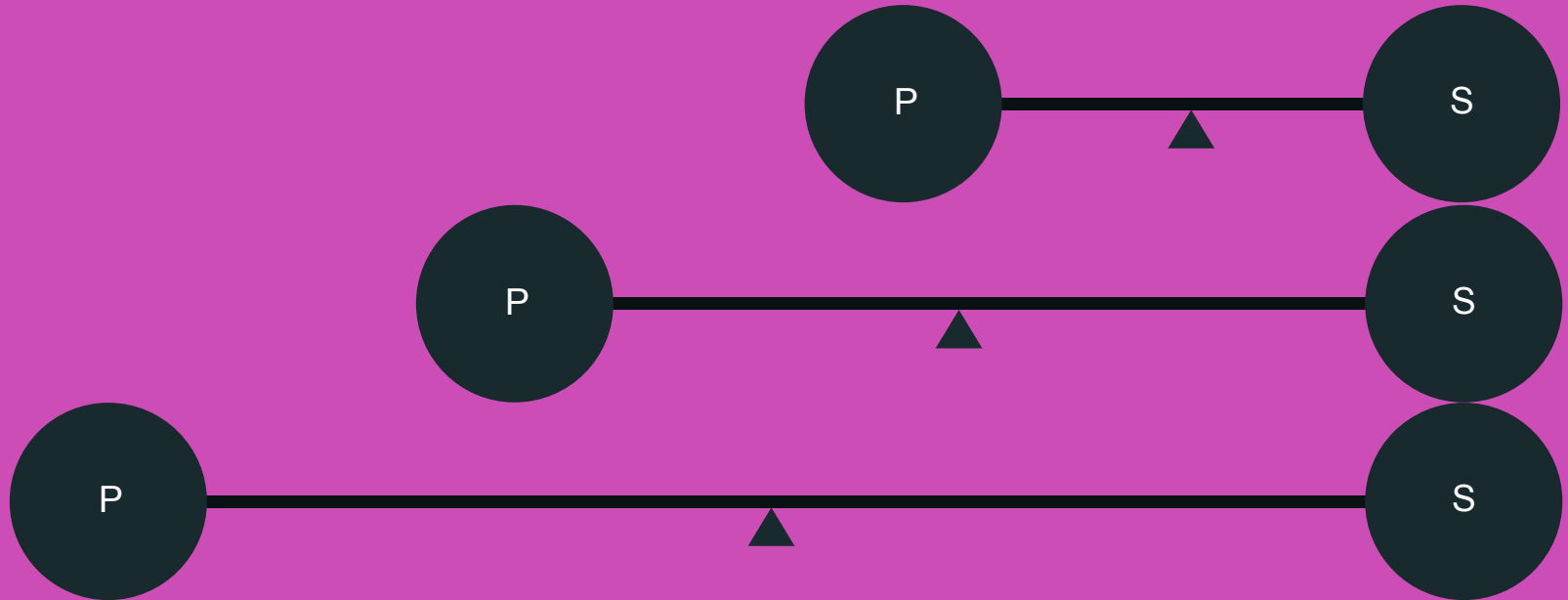
Centre of Mass



The **closer** the center of mass is to the star, the smaller the orbit of the star



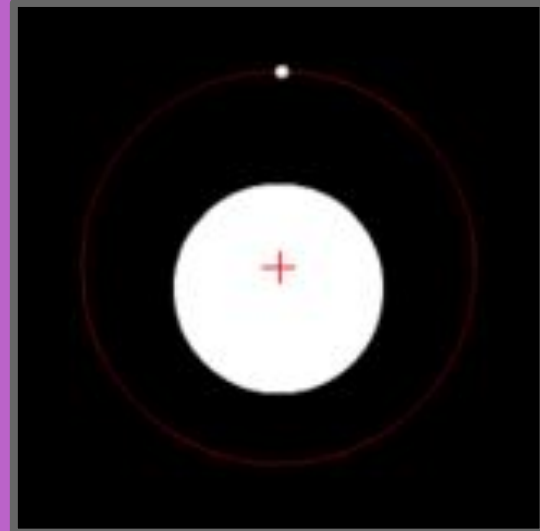
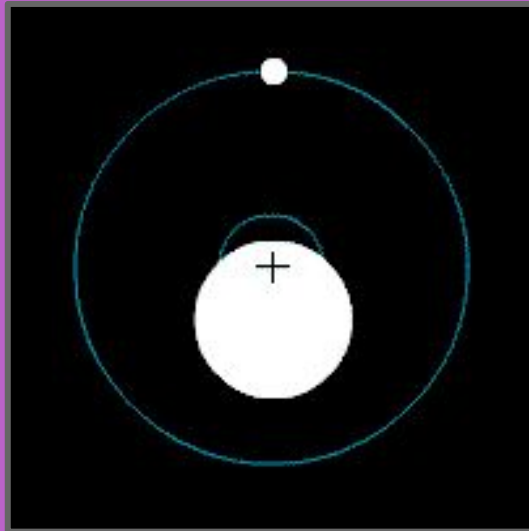
Centre of Mass



STARS DANCE



Even more when the planet is **HEAVY** and more rapidly if the planet is **CLOSE**.

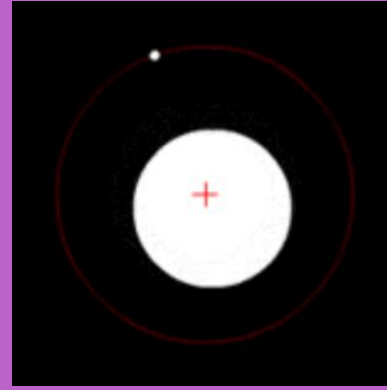
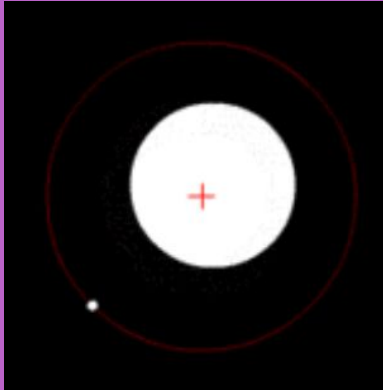




ASTROMETRY

Or the search for stars' invisible dance partners (part 1)

Definition: Precise measurement of location of objects in the sky.





ASTROMETRY

Or the search for stars' invisible dance partners (part 1)

1 exoplanet
found via
astrometry....

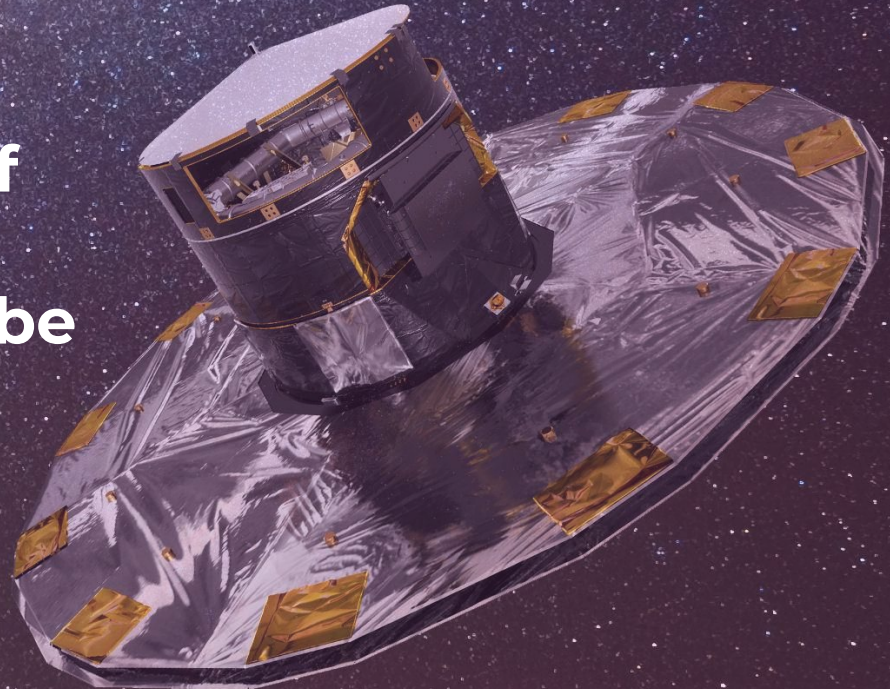




GAIA Mission

Actively searching for stars' invisible dance partners

10 of 1000s of
exoplanets
expected to be
discovered





ASTROMETRY

SUN



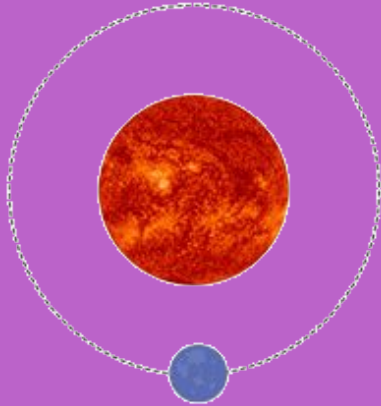
But these guys were found via astrometry!!!



Orbital Inclination

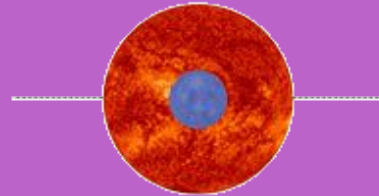
Face-On

$i=0^\circ$



Edge-On

$i=90^\circ$





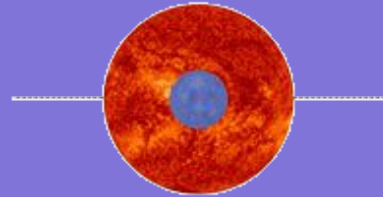
Radial Velocity

Or the search for stars' invisible dance partners (part 2)

Definition:

Precise measurement
the velocity of an
object towards and
away from you

$i=90^\circ$





Doppler Effect

With SOUND

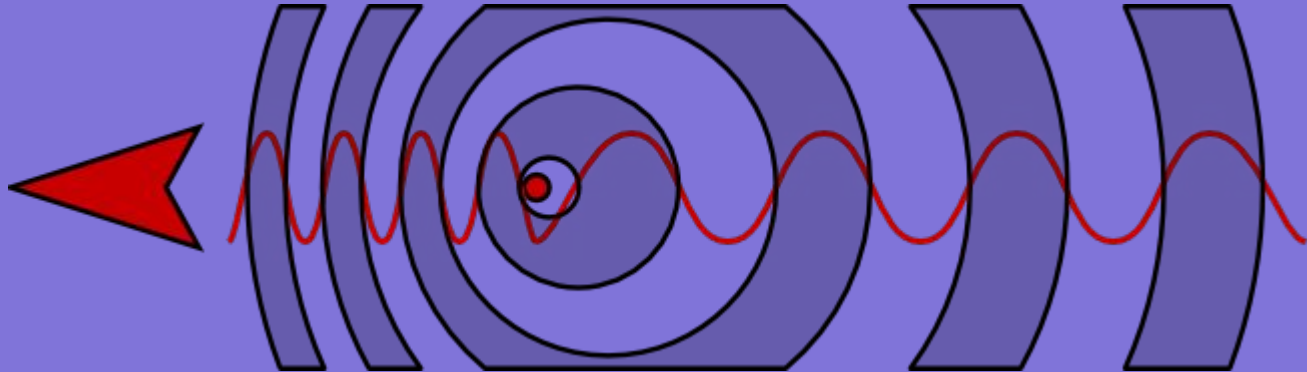
Higher
Frequency

Higher
Pitch



Lower
Pitch

Lower
Frequency





Doppler Effect

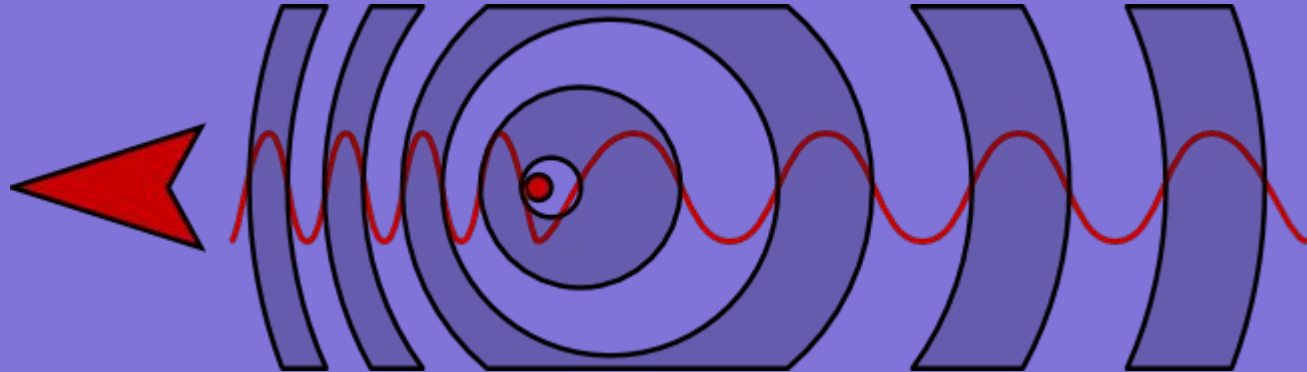
With Light

Higher
Frequency

BLUER

REDDER

Lower
Frequency





Doppler Effect

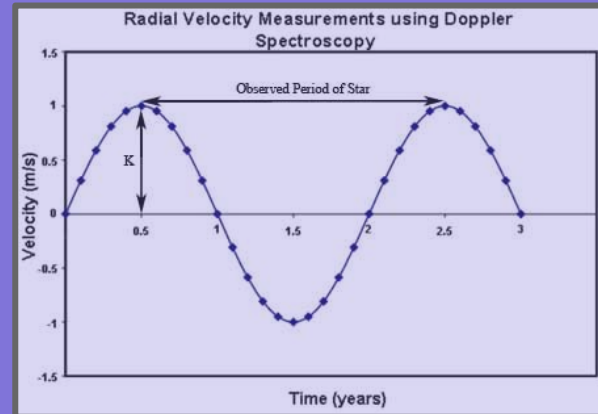
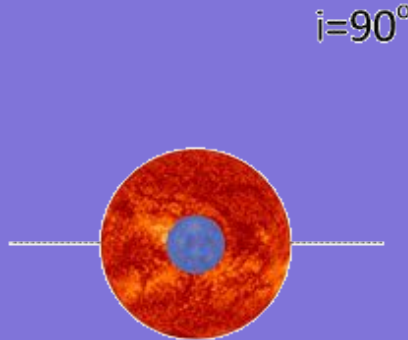
With Light

Higher
Frequency

BLUER

REDDER

Lower
Frequency



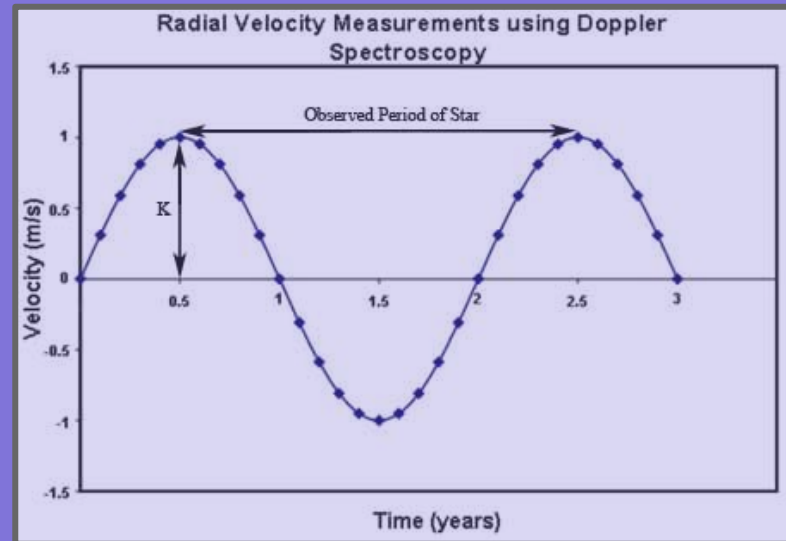


Radial Velocity

Or the search for stars' invisible dance partners (part 2)

ADDITIONAL INFO:

- Mass of the planet
- The size of its orbit
- eccentricity





Radial Velocity

Or the search for stars' invisible dance partners (part 2)

677 exoplanets
found via radial
velocity!



KECK OBSERVATORY

A hot Jupiter planet is shown in the foreground, appearing as a large, blue and orange sphere. In the background, a bright, yellow-orange star is visible, with a large, glowing ring of gas or dust around it. The background is a deep blue space filled with stars.

PLENTY

HOT JUPITER

WERE DISCOVERED!

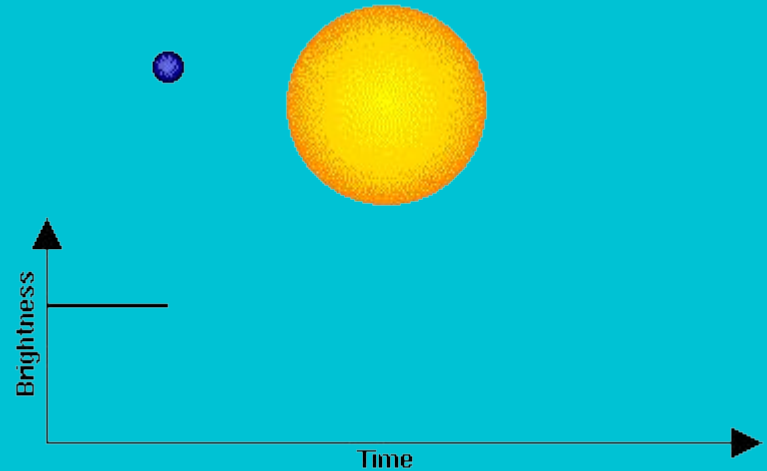


TRANSIT

or the search for shadows

When the planet **passes in front** of the star, it **blocks** incoming flux from the star.

The star then appear **dimmer** for a distant observer

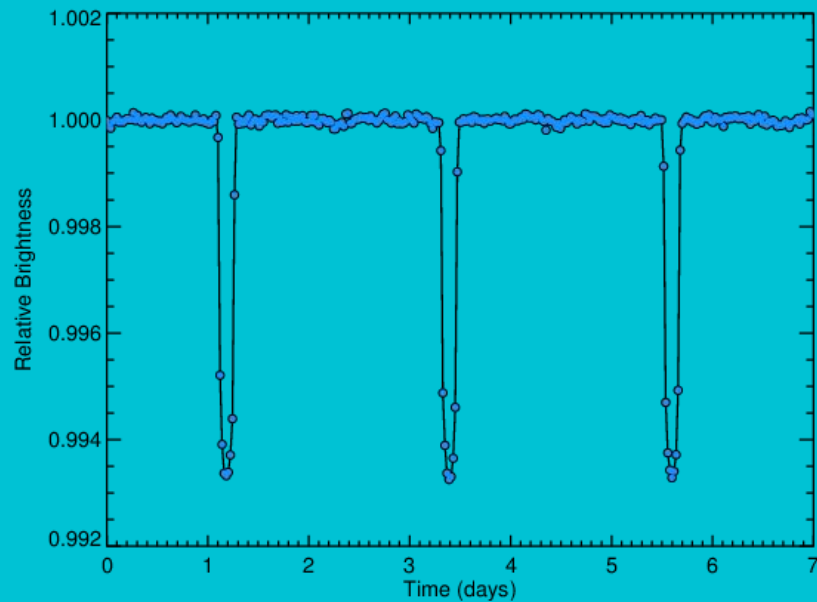




TRANSIT

HOW DO YOU KNOW
IT'S NOT JUST A
RANDOM OBJECT
PASSING IN FRONT?

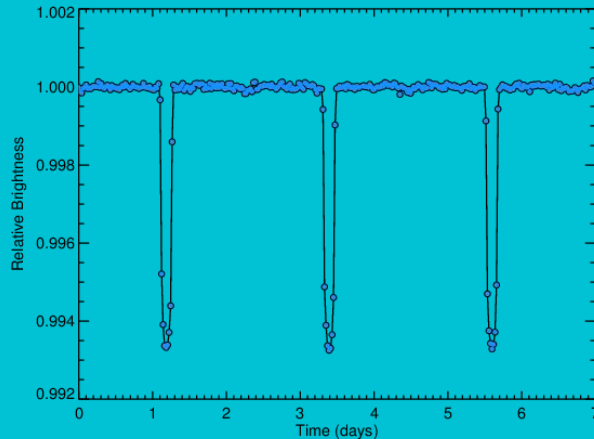
If it's an object in
orbit, it will pass in
front **again**.





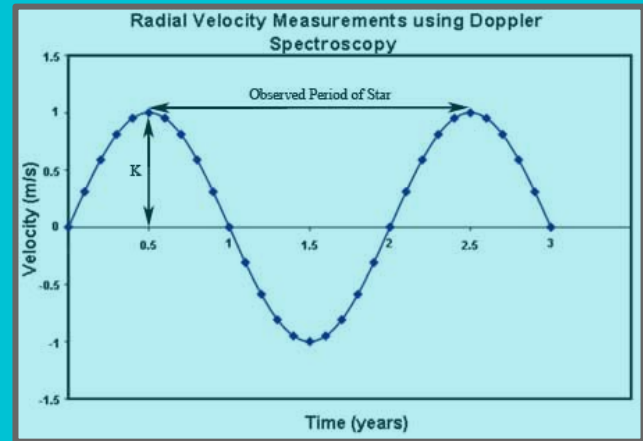
TRANSIT

**BUT, IS IT A
PLANET?**



We can only confirm that it is a planet if we know its **mass**.

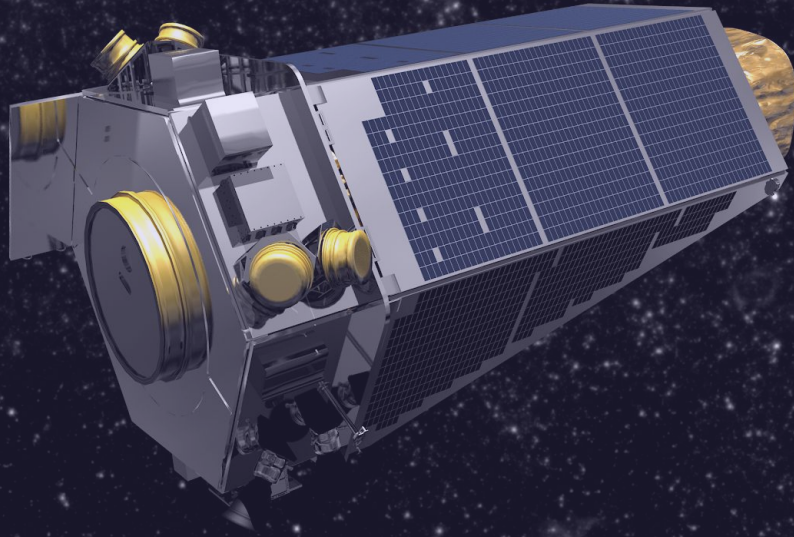
+





TRANSIT

Or the search for shadows



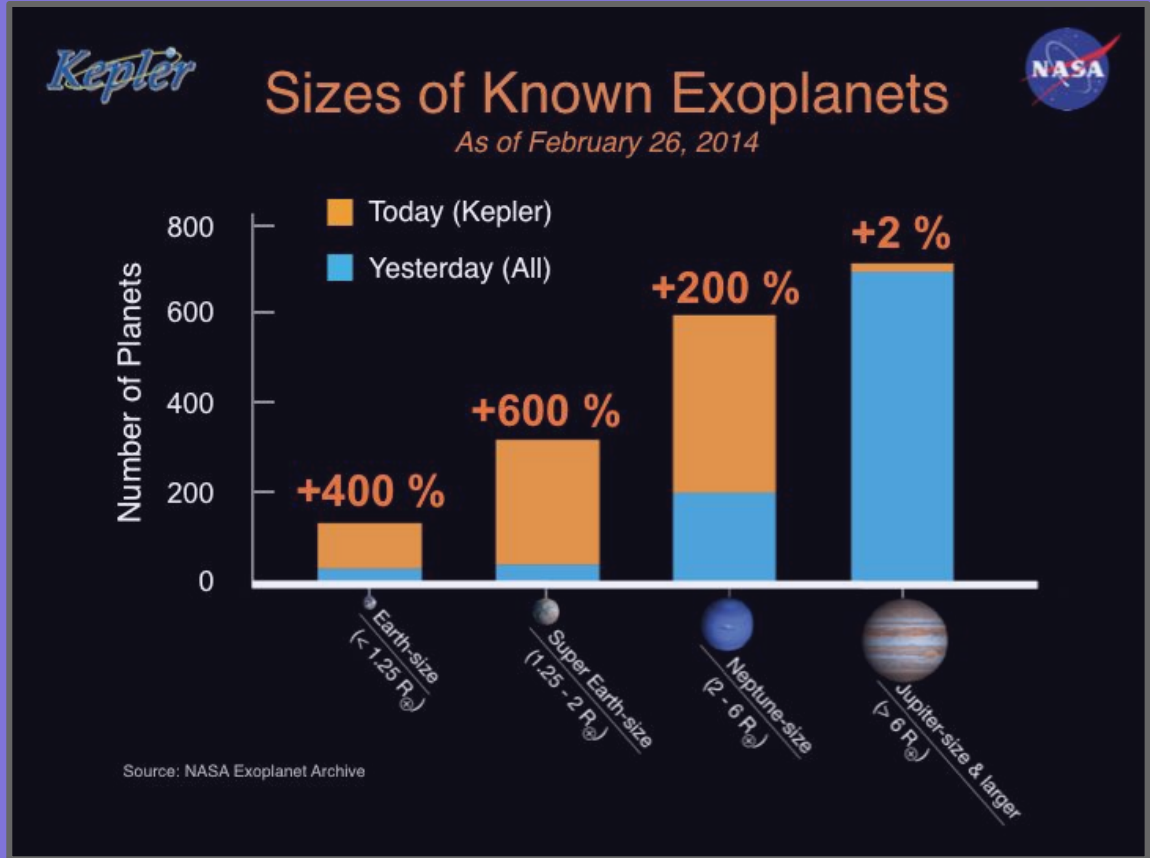
2955

exoplanets found via
transit! (and +4000 of
candidates!)

KEPLER SPACECRAFT



BEFORE KEPLER

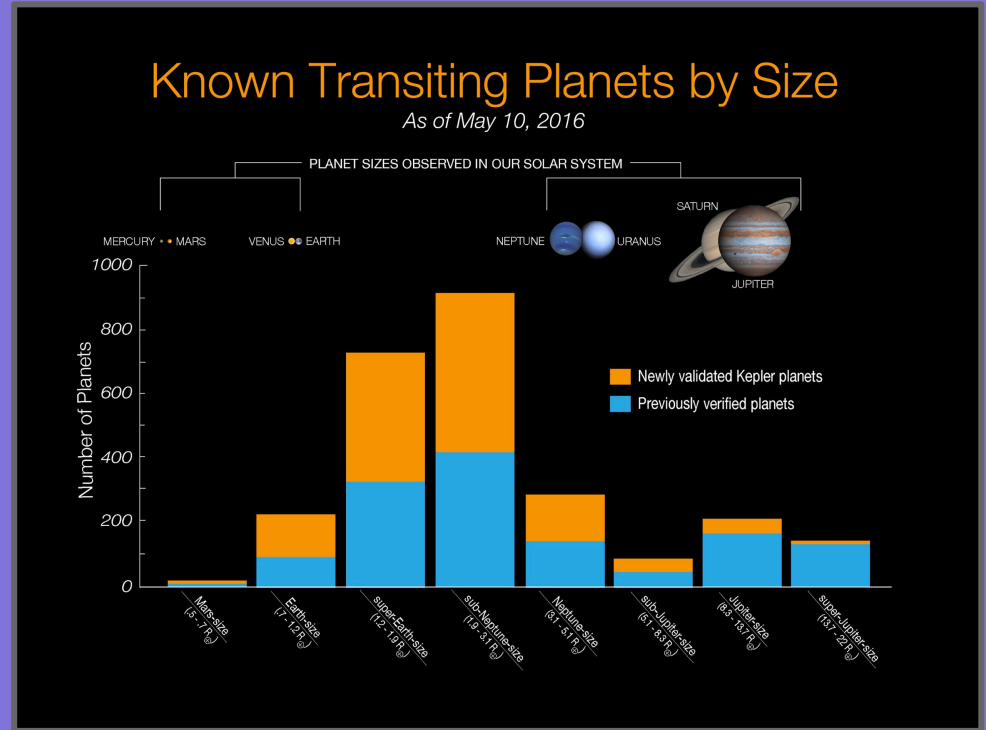


On May 10, 2016, the Kepler Team announced the discovery of



1284 EXOPLANETS

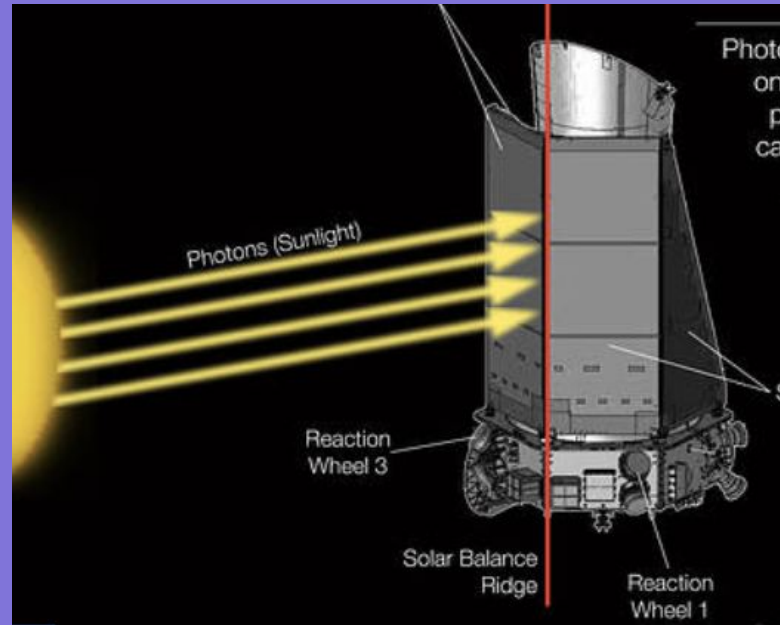
Sub-Neptune and Super Earth are the most common type of exoplanet





Kepler's Second Life: K2

In May 2013, the reaction wheels of Kepler failed

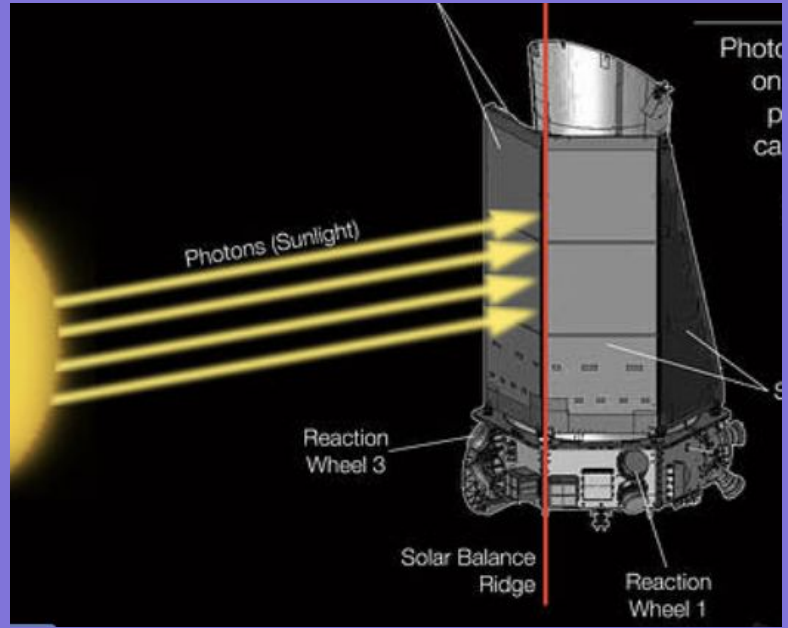




Kepler's Second Life: K2

325

exoplanets found via transit! (and +500 candidates!)



Kepler Search Space:
3000 light-years
0.25% of the sky



**Successfully launched
in April 2018**



TESS Search Space:
200 light-years
All-sky





HONORABLE MENTION

SPITZER

FIRST 7 PLANETS
SYSTEM FOUND!!!

TRAPPIST-1 SYSTEM



**TESS
FIRST
IMAGES!!!**

Tons of M-Dwarfs!!!!

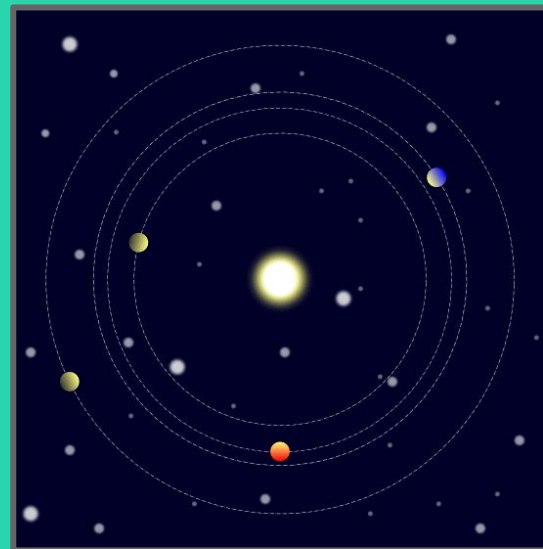


Transit Timing Variation

When 2 planets pass each other, they are **gravitationally attracted**.

They will **exchange energy**: one will speed up, the other will slow down.

The next transits will either occur slightly **before or after** expected.





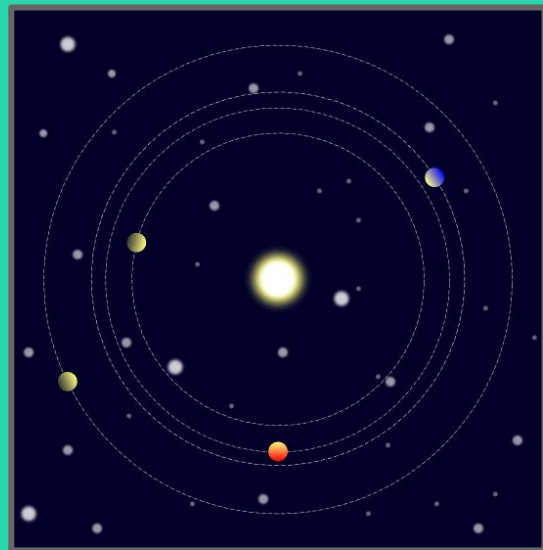


Transit Timing Variation

15

exoplanets
found through
TTV.

But more to
come?





Microensing

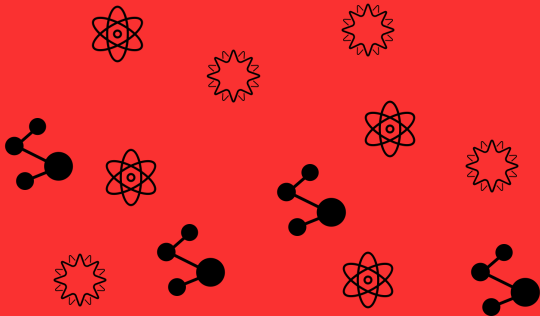
“Wait, so you are trying to find planets you can’t see around stars you can’t see?”

Dark Matter

5x more abundant than normal matter

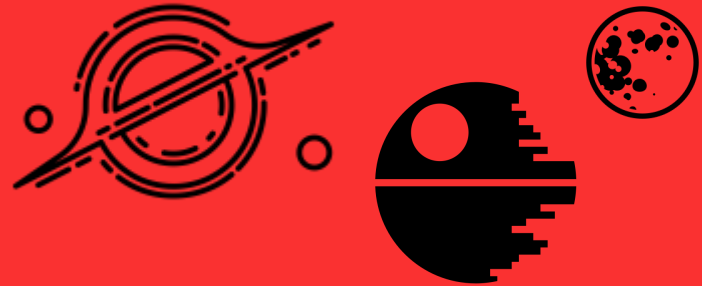
WIMPs

*(Weakly Interacting Massive
Particles)*



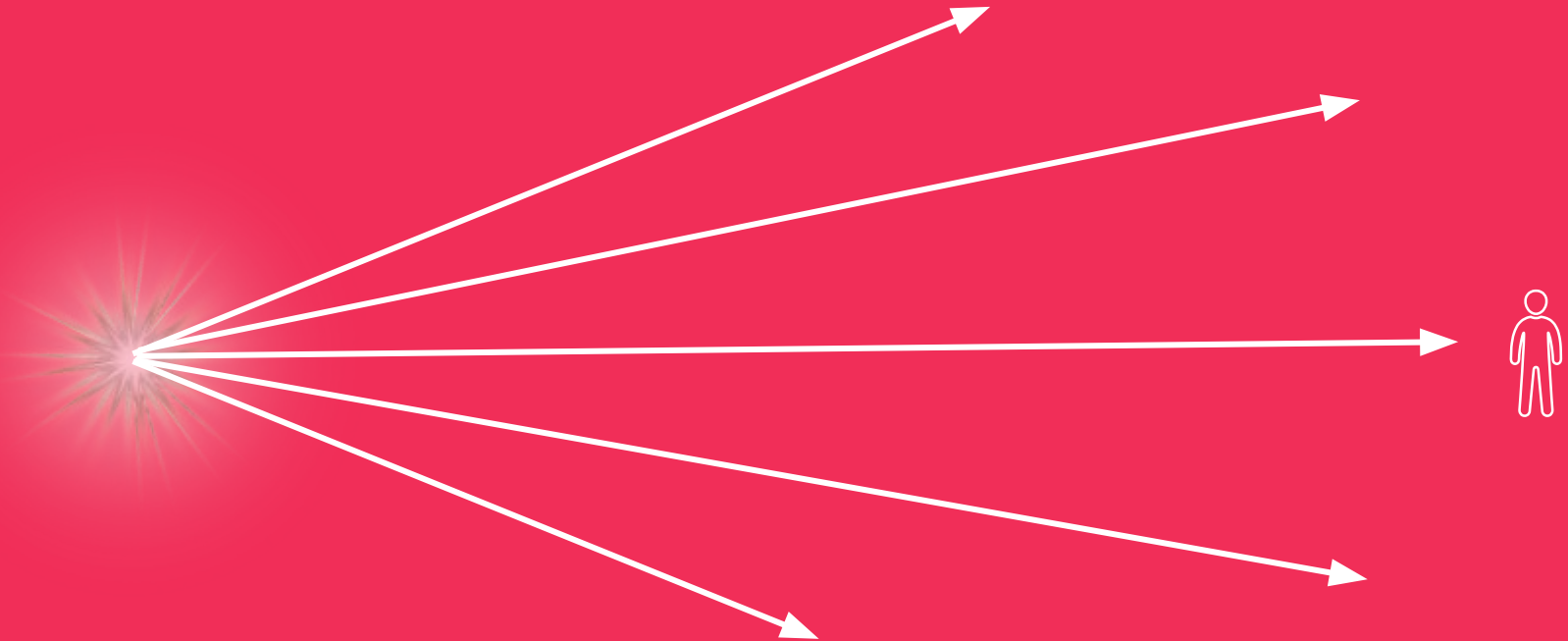
MACHOs

*(MAssive Compact Halo
Objects)*



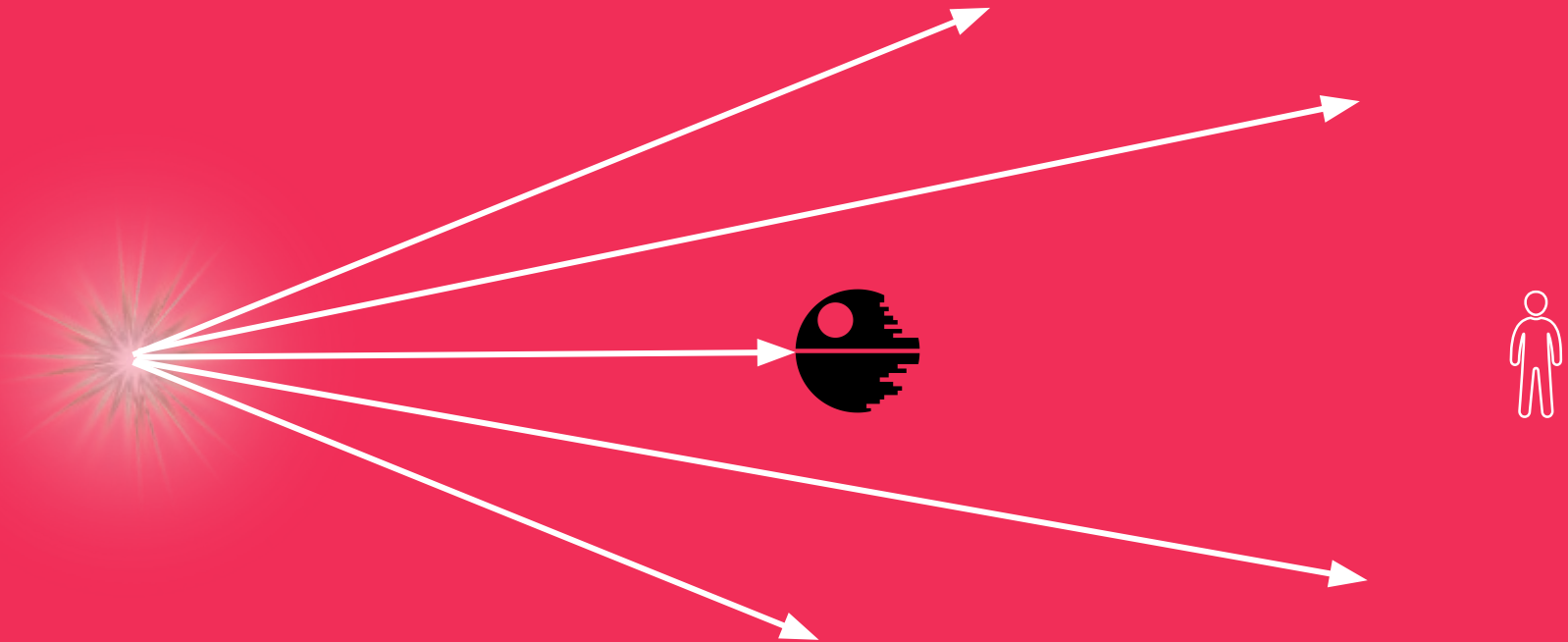
Trouver des MACHOs

Gravitational microlensing



Trouver des MACHOs

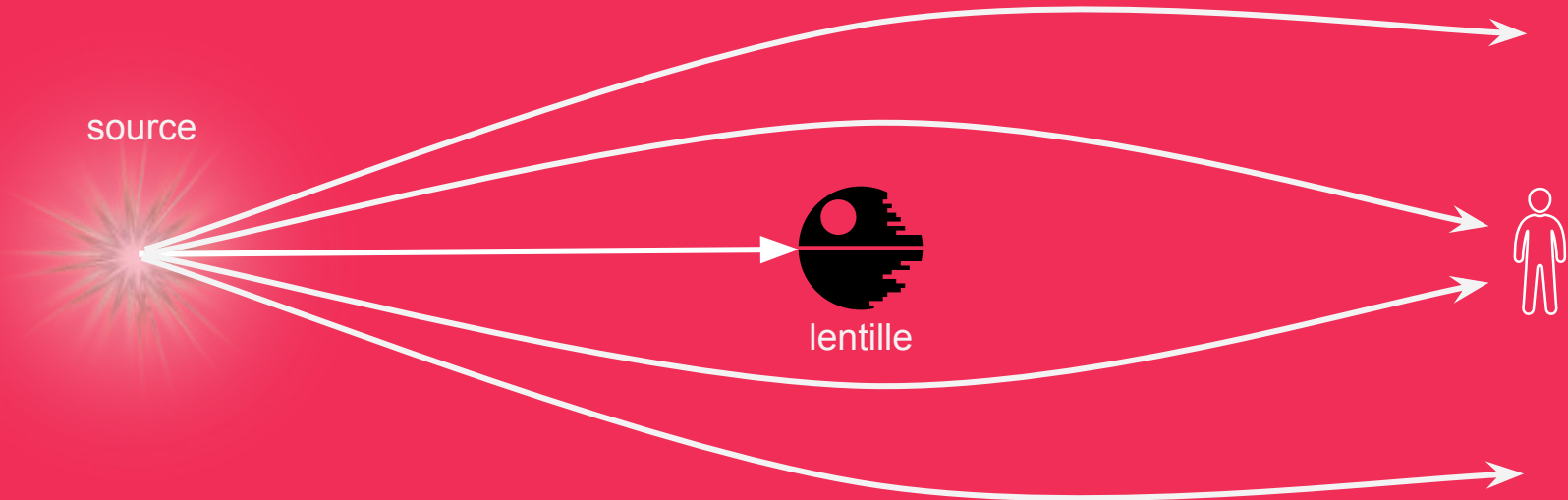
Microlentilles gravitationnelles



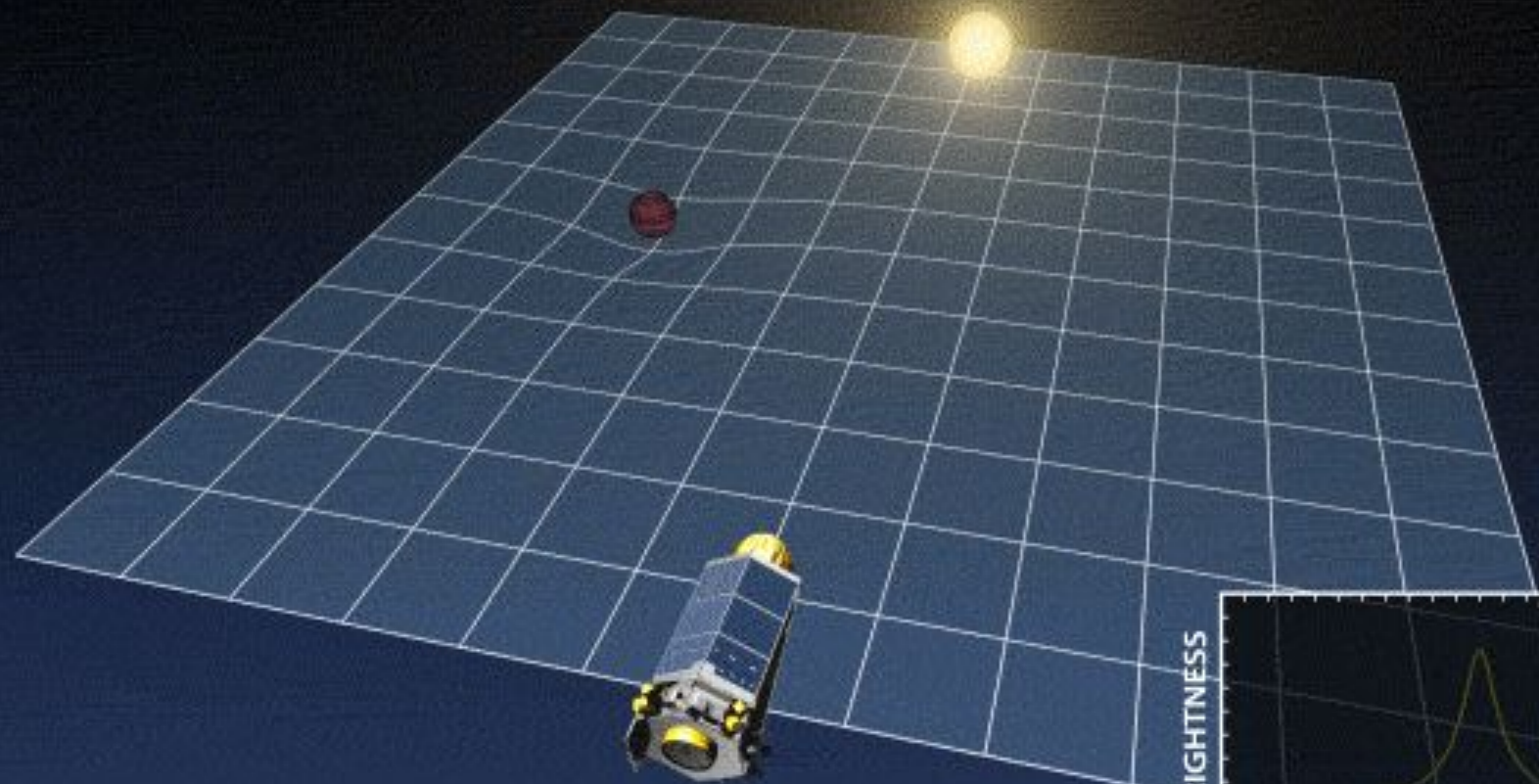
Finding MACHOs

Microlensing

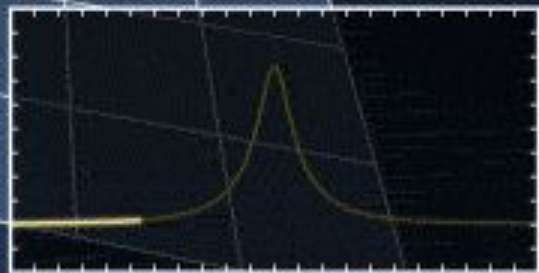
But Einstein said that mass bends the path of light!



So the mass in between **acts like a magnifier!**



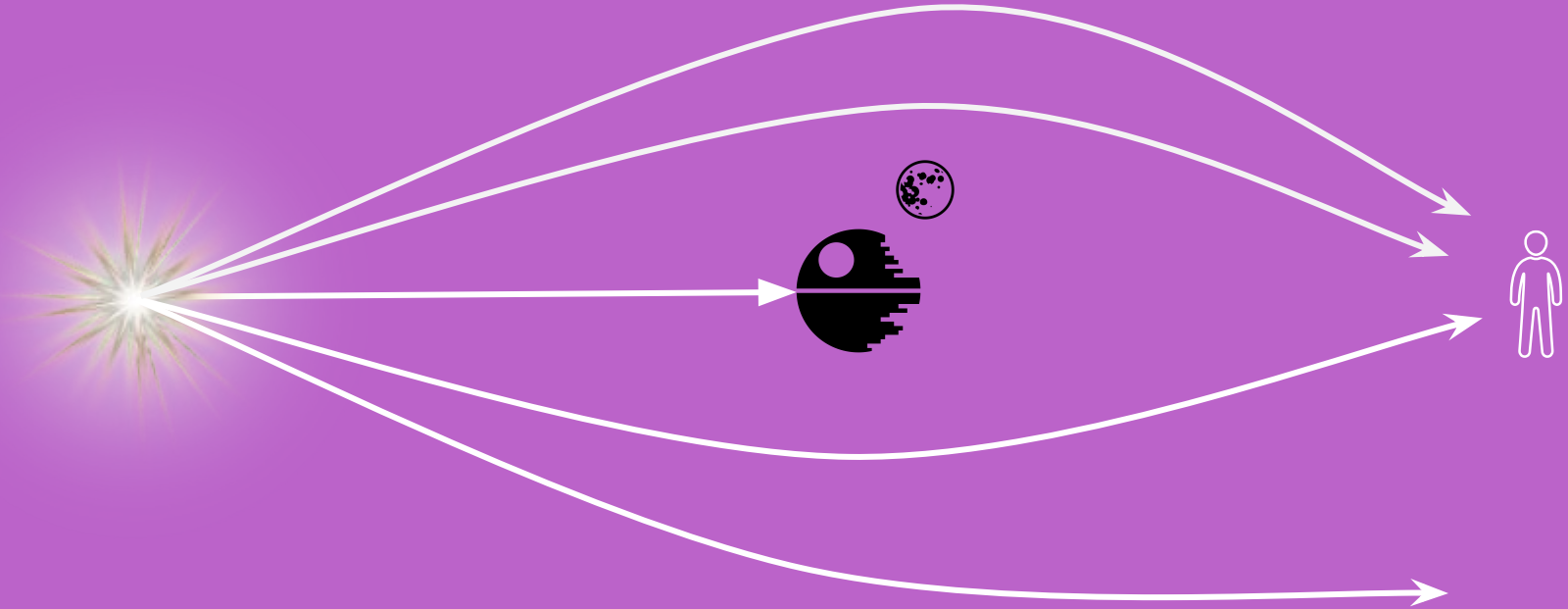
BRIGHTNESS



TIME

MACHOs

With a companion

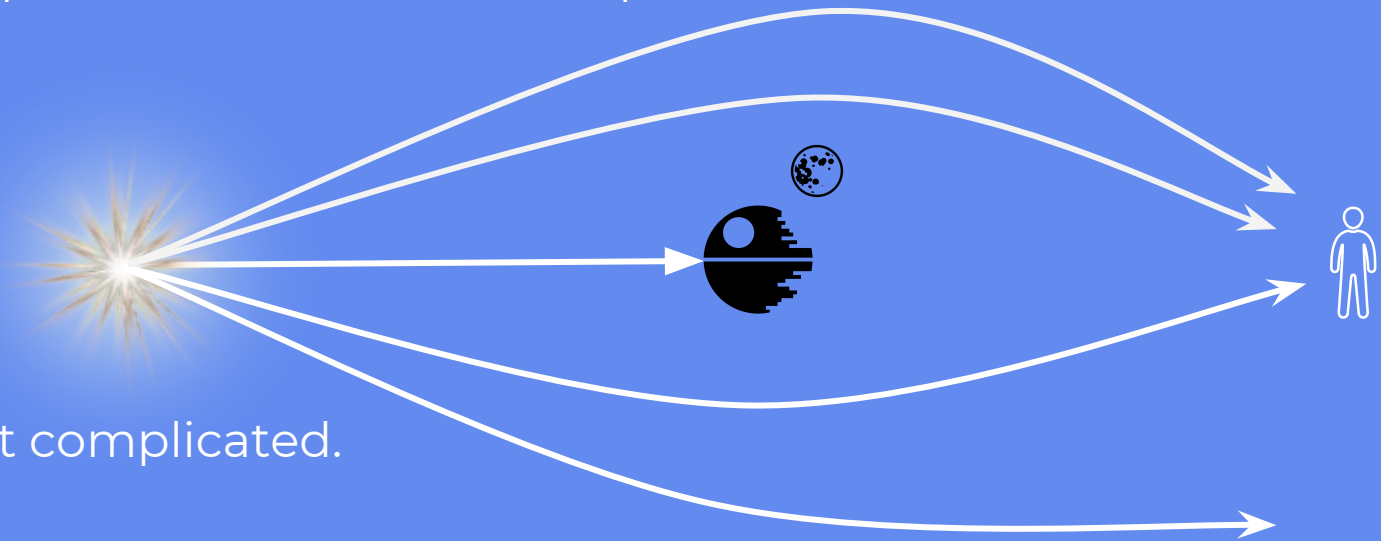




Microlensing

Aka the study of the lens.

What happens when there is a companion to the lens?



Things get complicated.

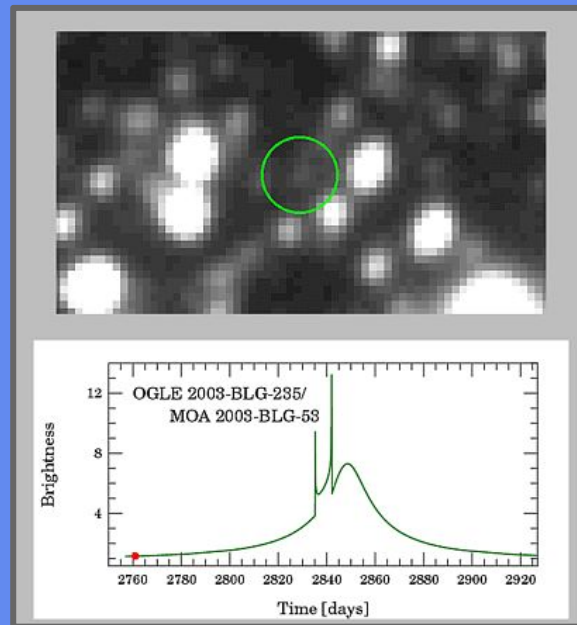


Microlensing

Aka the study of the lens.

The planet will cause the source to rapidly flash.

These flashes are unpredictable and last about a **couple of days.**





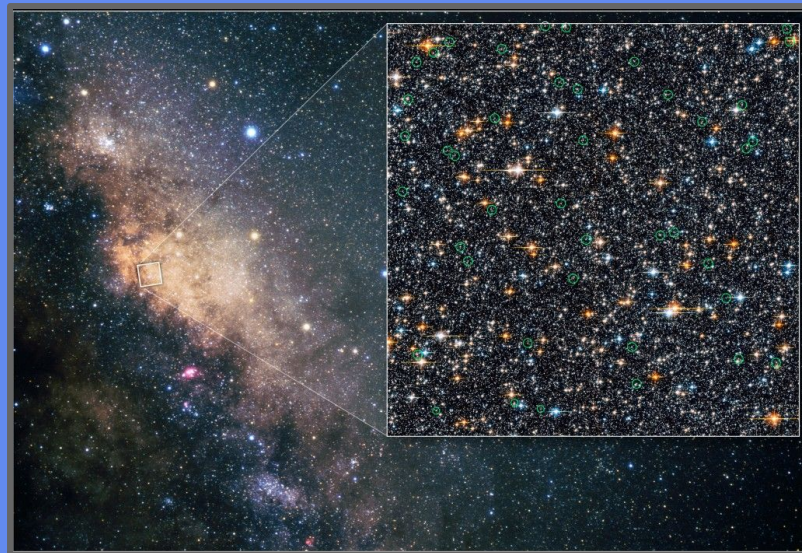
Microlensing

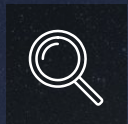
Aka the study of the lens.

Microlensing event are **unpredictable**.

If you stare at a star, this event will happen **once every million years**.

The strategy is to stare a 100 billions stars and hope for the best.





Microlensing

Procedure

1. Have a telescope looking at the bulge of the Milky way
2. Whenever a **star starts to get brighter**, put it on the public alert list.

If it gets **REALLY BRIGHT REAL QUICK**, make calls, tweet about it, alert everyone, stay up all night.



Microlensing

Procedure

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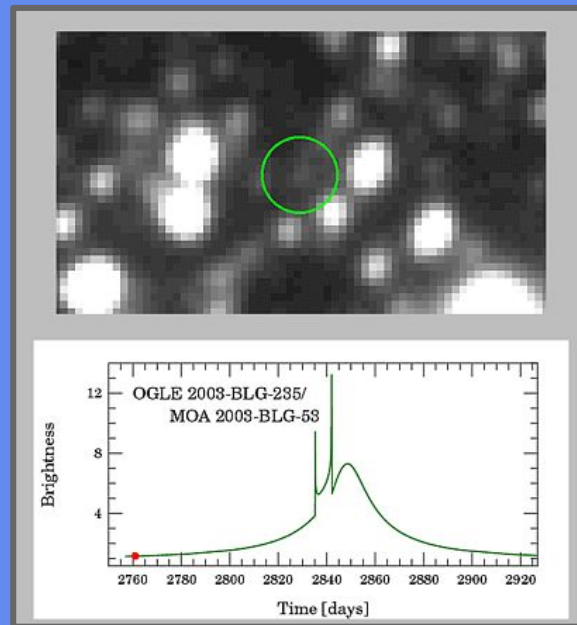
Microlensing

Aka the study of the lens.

64

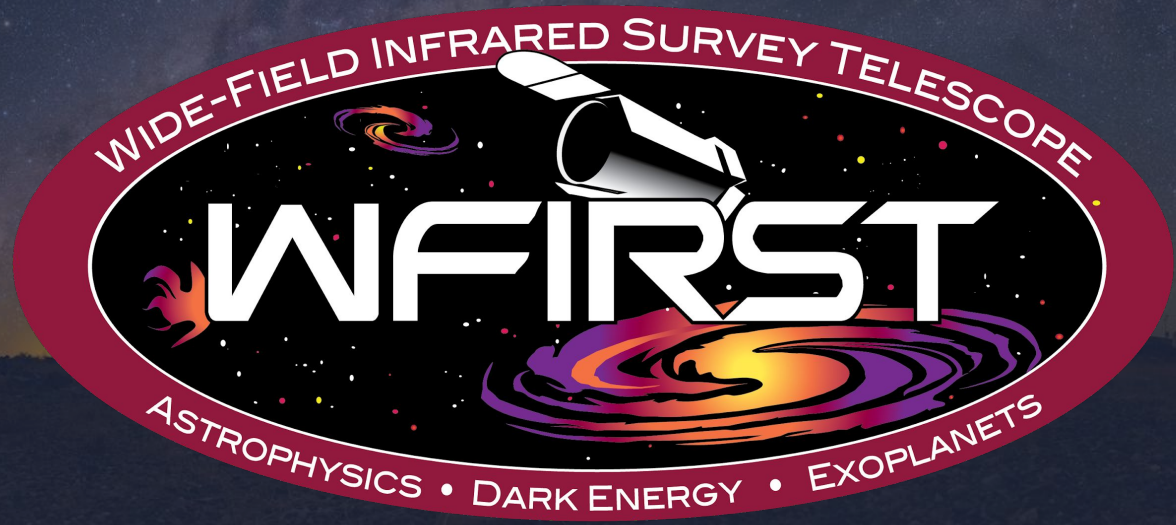
exoplanets
found through
microlensing.

Advantage,
finds planets on
wide orbit!!!





WFIRST





Direct Imaging

Definition:

Taking a picture of the planet. Duh-uh



Direct Imaging

IS HARD.

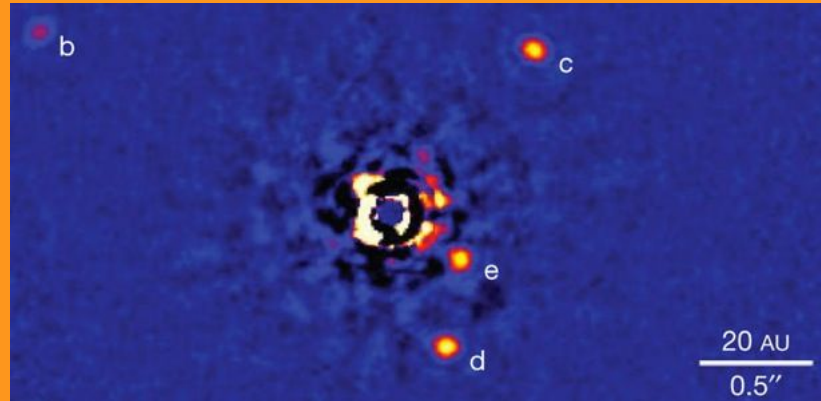
In fact, it is as difficult as trying to photograph a fly next to a lighthouse beam while standing over 6 kilometers away....



CORONAGRAPH

Definition:

Technique to separate the light from the star and then block the light from the star.

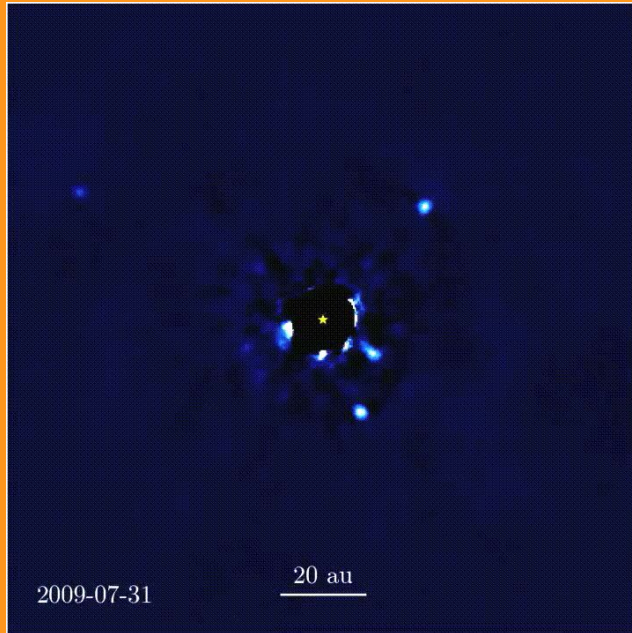




CORONAGRAPH

Definition:

Technique to separate the light from the star and then block the light from the star.





Direct Imaging

IS HARD, but not impossible

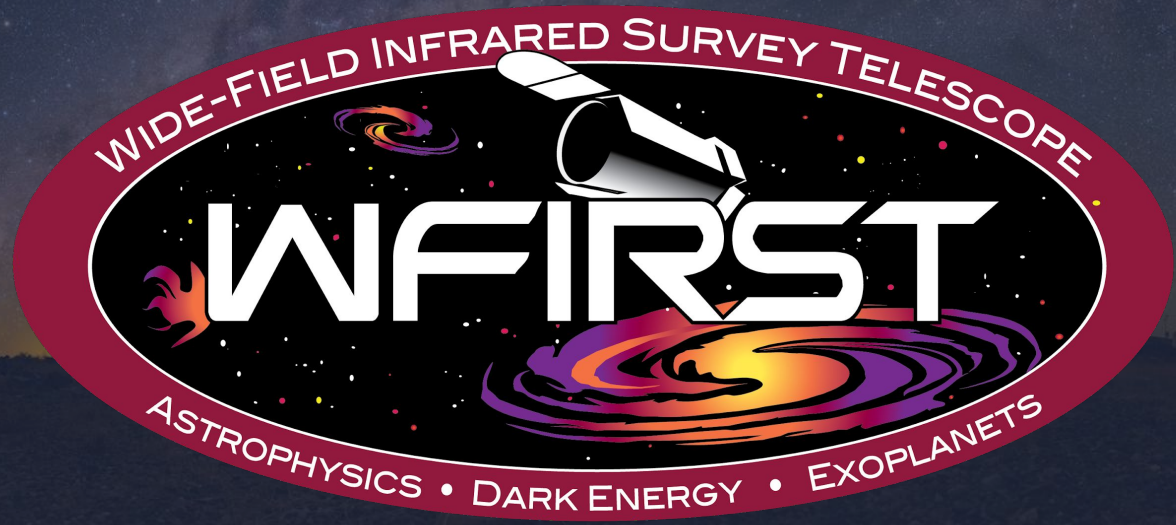
44

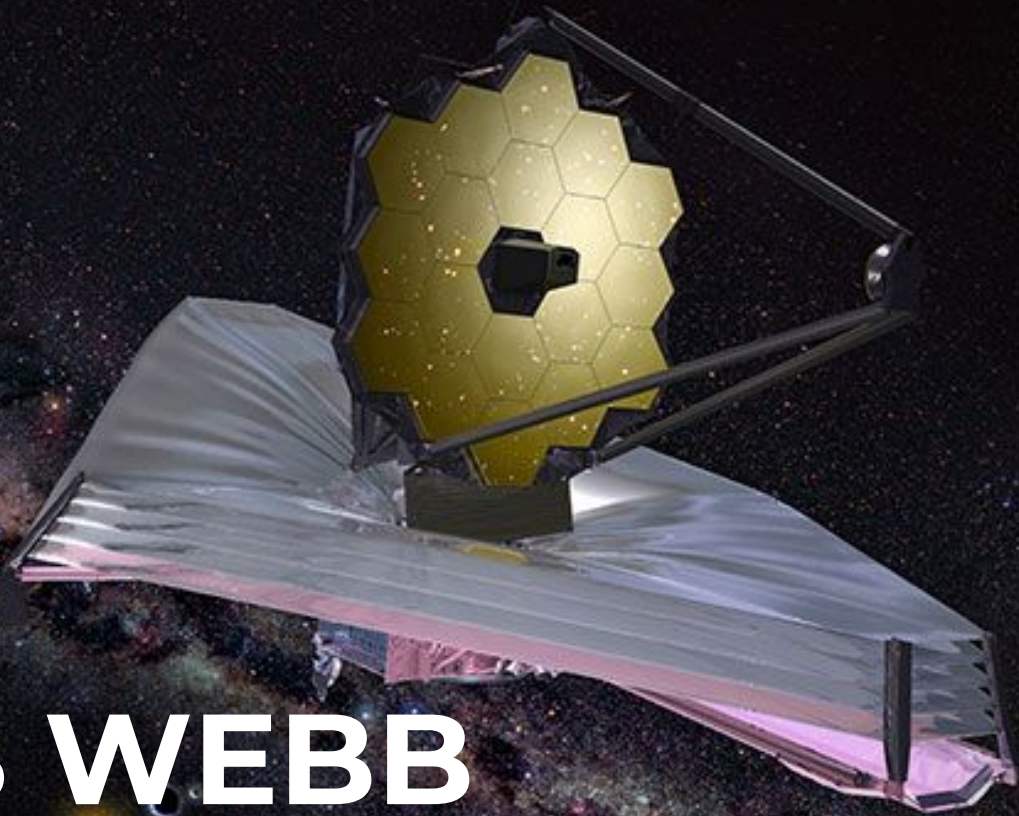
exoplanets
found through
direct imaging.

But more to
come?



WFIRST





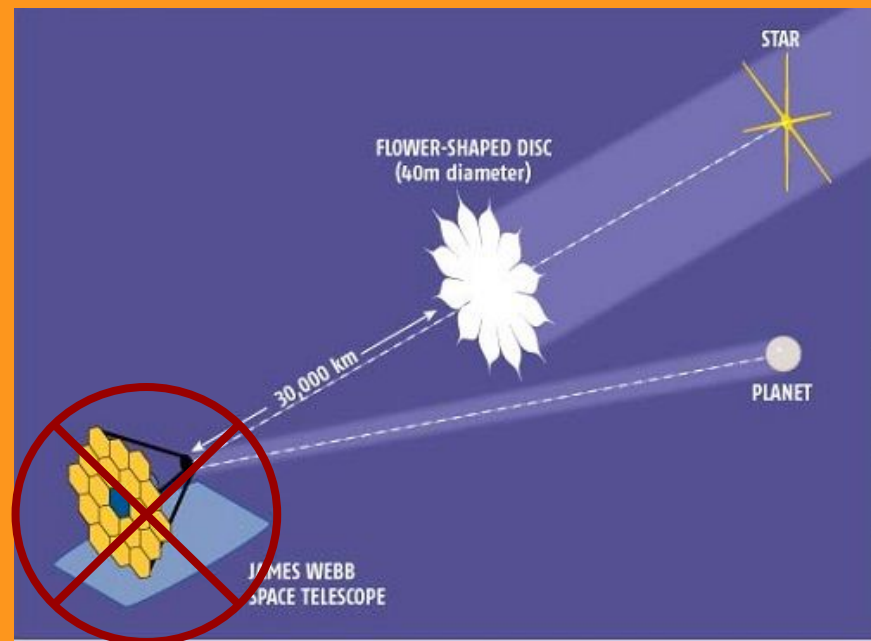
JAMES WEBB SPACE TELESCOPE

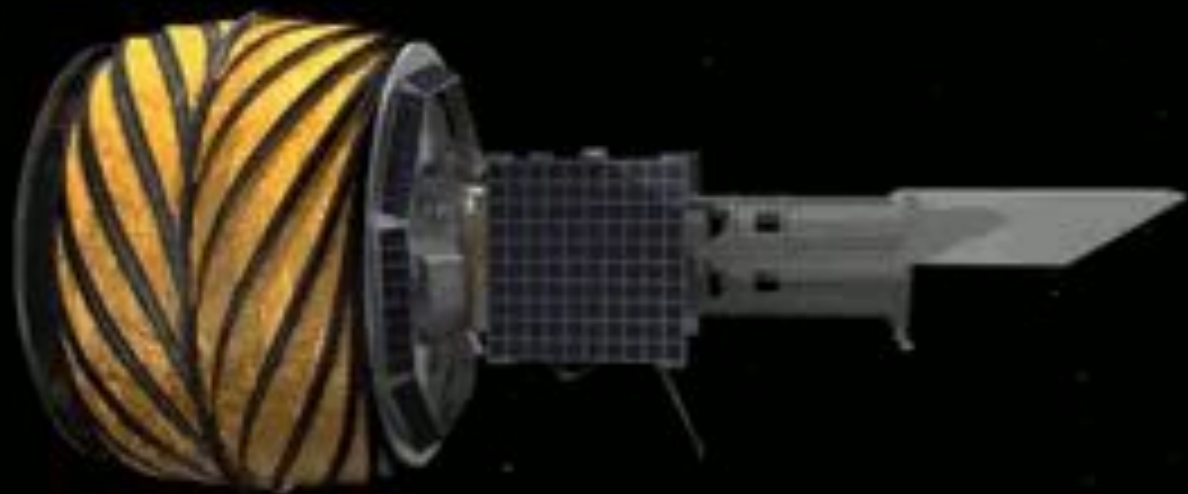


STAR SHADE

SPACE ORIGAMI

Blocking the light
from the star
externally



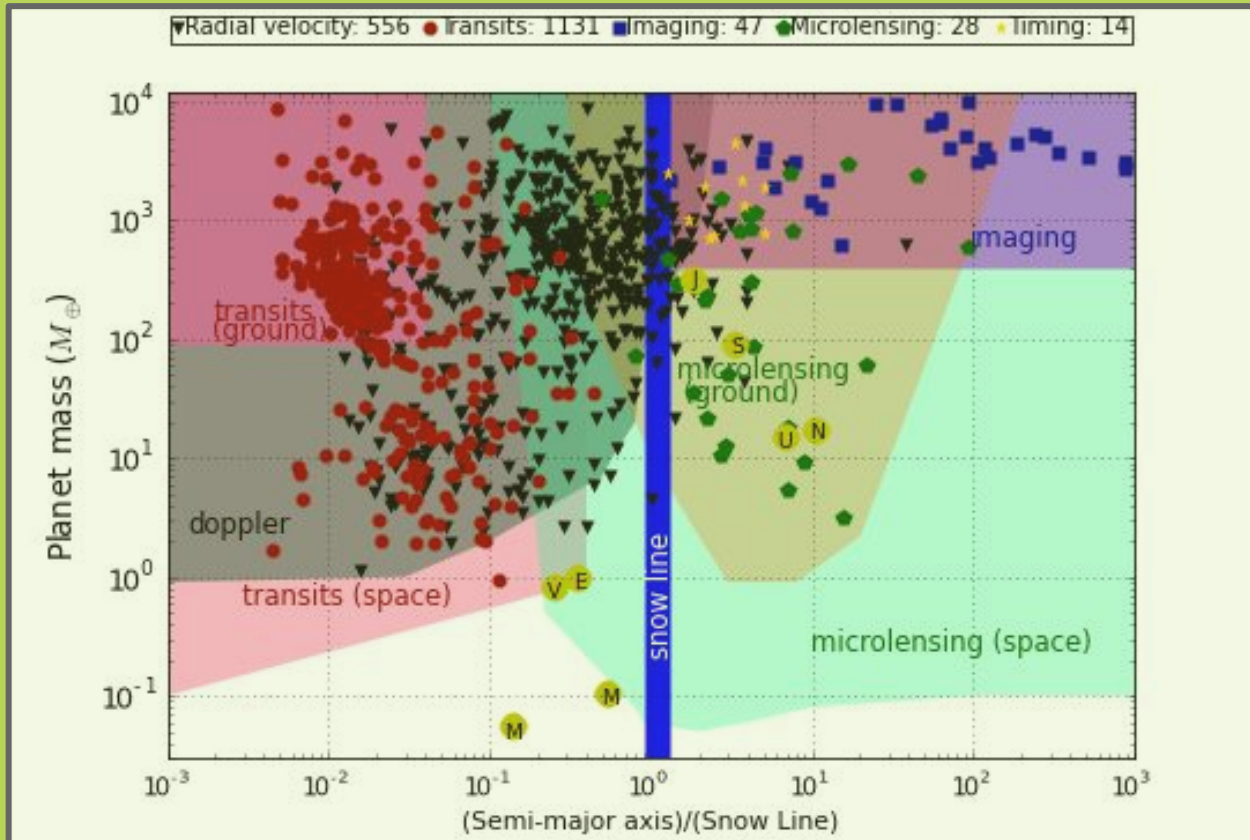




6 Methods to Detect Exoplanets

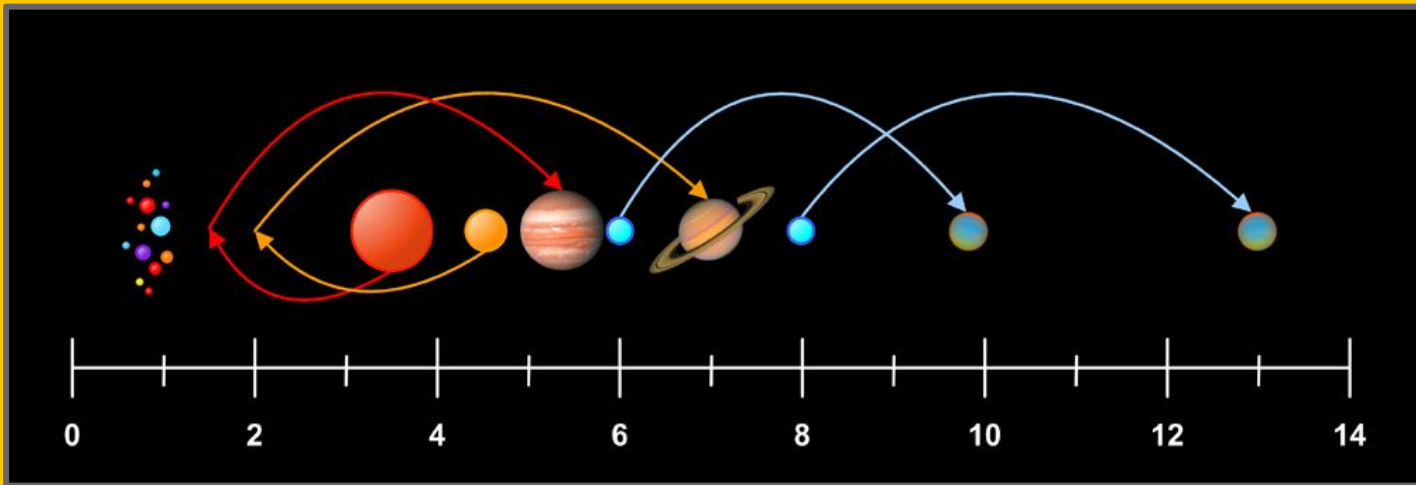
- ASTROMETRY (minuscule dance movements)
- RADIAL VELOCITY (periodically becomes bluer and redder)
- TRANSITS (blocking light from the star)
- TRANSIT TIMING VARIATION (non-equal period)
- MICROLENSING (taking advantage of general relativity)
- Direct Imaging (snapping photos)

WHY DO WE KEEP SEARCHING?



PLANET FORMATION?

OR THE PAST AND FATE OF THE SOLAR SYSTEM.





CLIMATE STUDIES

NOT TO BE CONFUSED WITH WEATHER...





ALIEN LIFE?

ARE WE ALONE
IN THE UNIVERSE?



PLANET HOP FROM
TRAPPIST-1e



VOTED BEST "HAB ZONE" VACATION WITHIN 12 PARSECS OF EARTH



**NOT A
SEARCH
FOR A
SECOND
HOME**

A large radio telescope dish is suspended from a complex metal structure against a sunset sky. The dish is dark and hangs from a network of cables and beams. The sky is a mix of orange, yellow, and blue, with some clouds. In the background, a tall, thin tower is visible. The overall scene is silhouetted against the bright sunset.

Public AstroNight

*Pinging space rocks
at the Arecibo Observatory*

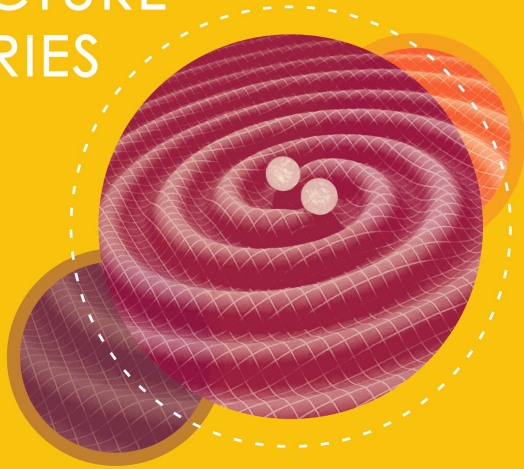
Dr. Flaviane Venditti

from Arecibo Observatory

Thursday, Sept. 20

**Tonight
7 PM
McConnell
Room 204**

PUBLIC
LECTURE
SERIES



UNPACKING GRAVITATIONAL WAVES:
THE ROAD TO DISCOVERY

TALK BY BENJAMIN DRINGOLI
From McGill University

Location: Arts Building W-215
Metro McGill

Thursday, September 27, 7 PM

More info: physicsmatters.physics.mcgill.ca/lectures/



**Next
Thursday
7pm**

**Arts Build.
Room W-215**